

1 **Wavelet multi-resolution analysis of recent** 2 **GOCE/GRACE GGMs**

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7 **Abstract**

8 The realization of the GRACE/GOCE missions offer new opportunities for gravity field
9 approximation with higher accuracy at the medium wavebands, while wavelets (WL) provide
10 powerful gravity field analysis tools in the space/frequency domain. This work focuses on the
11 spectral analysis of GOCE, GOCE/GRACE and combined Global Geopotential Models (GGMs)
12 through wavelet decomposition, filtering and reconstruction to improve their performance in the
13 higher bands of the spectrum. The GGMs evaluated refer to the latest DIR-R4/R5, TIM-R4/R5 and
14 GOCO03s models, which are compared with local GPS/Leveling geoid heights and gravity
15 anomalies, while EGM2008 is used as a reference. Through a WL-based multi-resolution analysis,
16 gravity anomalies and geoid heights are analyzed to derive their approximation and detail
17 coefficients for various levels of decomposition, which correspond to different spatial scales. The
18 content and signal power of each level of decomposition is analyzed to conclude on the amount
19 and quality of signal power that GOCE/GRACE GGMs represent compared to EGM2008,
20 especially up to the targeted waveband of 100-150 km. Filtering is investigated as well to remove
21 high-frequency information from the low resolution GOCE models and adjust the WL
22 reconstruction. The model synthesis that follows, through WL coefficient reconstruction, aims at
23 the generation of new synthesized GGMs, where both GOCE and EGM2008 information is used,
24 the latter serving to model the omission error in the GOCE GGMs. The synthesized GOCE GGMs
25 offer an improvement of more than 30 cm compared to the original GOCE GGMs, while they
26 provide a 1-2 cm improvement compared to EGM2008. In terms of the validation with gravity
27 anomalies, a 5 mGal improvement was found, w.r.t. to the original GOCE GGMs, while w.r.t.
28 EGM2008 there was no improvement. Finally, it was concluded that the GOCE GGMs show
29 improved, between 5-22%, correlation with the land topography compared to EGM2008 for spatial
30 scales between 176-704 km.

31

32

33 **Keywords**

34 Wavelets, gravity field, multi-resolution analysis, filtering, spectrum coherency,
35 correlation, validation.

36 **1. Introduction**

37 Monitoring and understanding of the Earth's gravity field parameters at various
38 spatial scales has been the focus of many studies during the past decades. The
39 satellite missions of GOCE and GRACE have provided new insights to gravity
40 field monitoring and interpretation. Moreover, significant results related to the
41 time variation and evolution of the gravity field have emerged, the latter being a
42 result of mass/water redistribution in system Earth as well as a response to
43 geodynamic phenomena e.g., mega-earthquakes (Fuchs et al., 2013). GOCE has
44 offered invaluable data on sea-level change, ocean currents and circulation and ice
45 dynamics (Knudsen et al., 2011; Menna et al., 2014). Especially as far as gravity
46 field and geoid approximation is concerned, GOCE has offered improved,
47 compared to EGM2008, representations of the medium band of the spectrum, i.e.,
48 degree and order between 160 and 220, by as much as 4-5 cm over Europe
49 (Gruber et al., 2011; Hirt et al., 2011; Vergos et al, 2014).

50
51 This work focuses on the evaluation of the spectral content of GOCE/GRACE-
52 based GGMs, both satellite only and combined ones, by assessing their accuracy
53 in terms of both gravity anomalies and geoid heights. Gravity anomaly evaluation
54 is carried out through local gravity measurements covering the entire European
55 continent. Moreover, an extensive network of collocated GPS/Leveling
56 benchmarks, covering continental Greece, are used for the geoid height
57 evaluation.

58
59 Contrary to the usual evaluation in terms of the Global Geopotential Model
60 (GGM) combination with EGM2008 (see e.g., Gruber et al., 2011; Vergos et al.,
61 2014), in this work we employ wavelet (WL) decomposition as a multi-resolution
62 analysis (MRA) tool. MRA describes the (infinite) sequence of closed linear
63 subspaces of the space of square integrable functions $V_j \subset L^2(\mathfrak{R})$ and can be
64 applied with or without data. The transition from a MRA model $\{V_j\}$ to an
65 associated wavelet basis function is what constitutes the decomposition scheme
66 for every MRA subspace. Although WL MRA is relatively new as an analysis tool
67 compared to Fourier analysis, wavelets have been developed in order to overcome
68 the deficiencies of the Fourier transform (Mallat, 1989). The advent of the WL

69 transform and WL modelling in geosciences brought flexibility in the analysis
70 process for over a decade, since it allows the decomposition of the signal under
71 study to specific spatial scales that correspond to the levels of WL decomposition.
72 This is especially important since it allows the study of the properties of each
73 individual level (corresponding to specific spatial scales) without influencing the
74 rest. Within that frame, WL have been employed for local/regional determination
75 of the Earth's gravity field (Panet et al., 2011), the identification of large-scale
76 geoid undulations and their relation to mantle processes (Hayan et al., 2012), the
77 solution of the Altimetry-Gravimetry Boundary Value Problem (Grebenitcharsky
78 and Sideris, 2005) and lately to the analysis of GOCE satellite gravity
79 gradiometry data (Grebenitcharsky and Moore, 2014). In this work WL transform
80 and MRA are used to analyse both gravity anomalies and geoid heights in
81 approximation and detail coefficients for various levels of decomposition. Given
82 the initial resolution of the available data, the decomposition levels can be
83 translated to spatial scales, hence WL MRA allows the study of GOCE/GRACE
84 GGM contribution to various parts of the gravity field spectrum. To improve the
85 performance of GOCE/GRACE GGMs, as to their spectral content in the higher
86 bands of the spectrum, they are combined through wavelet decomposition,
87 filtering and reconstruction with EGM2008. Hence EGM2008 is used to model
88 the omission error in the low-degree GOCE/GRACE GGMs. Both the original
89 and synthesized GOCE/GRACE fields are evaluated with in-situ gravity
90 anomalies and GPS/Leveling observations on benchmarks (BMs).

91

92 Coherence and correlation are powerful tools for testing the relation between land
93 topography and gravity anomalies (Δg) for different spectral bands. They are both
94 employed over the Amazon area and Europe, in combination with WL
95 decomposition, to determine the coherency and correlation of GOCE/GRACE
96 GGMs for different bands of the spectrum.

97

98 **2. Methodology, GGMs and data availability**

99 **2.1 GOCE/GRACE GGM analysis**

100 Wavelets are base functions with localization properties in both space (time) and
101 scale (frequency) domains and allow the synchronous analysis of GGM data at
102 different levels/scales. Therefore wavelet signal processing can be a
103 multiresolution analysis (MRA) at various levels of decomposition (Chui, 1992).
104 The two-dimensional wavelet transform gives coefficients that correspond to
105 different spatial resolutions, related to the signal scales (Grebentcharksky and
106 Moore, 2014). According to the WL decomposition algorithm, each scale (level
107 L_n) of the signal is analyzed in an approximation coefficient (A_n), which carries
108 the main (large scale) information and three detail coefficients (horizontal,
109 vertical and diagonal $(H, V, D)_i |_{i=1,2,\dots,n}$ that carry the high-frequency (short-
110 scale) information of the signal (Mallat, 1989; 1999). Through the synthesis
111 process, various GGMs can be combined, since each level can be represented by a
112 different GGM, based on the GGM performance at each specific level of analysis.
113 Of course, if after the WL MRA the WL coefficients remain unaltered during the
114 reconstruction process, the the original signal will be reconstructed perfectly due
115 to the orthogonality of the WL base functions.

116
117 The synthesis is defined as the algebraic sum of the detail coefficients of each
118 level used $(H, V, D)_i |_{i=1,2,\dots,n}$ and the approximation coefficient of the last level
119 (A_n) as:

$$120 \quad \textit{Synthesis} = A_n + (H, V, D)_n + (H, V, D)_{n-1} + \dots + (H, V, D)_1 \quad (1)$$

121 Therefore, for the available GOCE/GRACE GGMs, their spectral content at each
122 level is analyzed and compared among each other and with EGM2008. The goal is
123 to construct combined GGMs where both GOCE/GRACE and EGM2008
124 information will be used, so that the gravity field signal will be represented with
125 higher accuracy. The choice of the GGM that will be used at each level depends
126 on its resolution and the gravity field content with respect to EGM2008. Then,
127 different GGMs can be combined during the synthesis process in order to

128 determine a combined/hybrid GGM. In that sense, the first levels of the EGM2008
129 decomposed signal (small spatial scales) can be used to model the omission error
130 in the GOCE GGMs. Likewise, the decomposed GOCE GGM signal can provide
131 gravity information for those levels that correspond to the spatial scales targeted
132 by the GOCE mission (larger than 100 km). Hence, during the synthesis process,
133 various levels from different GGMs can be combined, in order to provide
134 improved representations of the Earth's gravity field. This aims at reducing the
135 omission error in GOCE GGMs and augmenting EGM2008 with gravity
136 information from the GOCE mission.

137

138 Moreover, when GOCE/GRACE GGMs are analyzed, the gravity signal of the
139 first levels (high-frequencies) is dominated by noise since these spatial scales are
140 not mapped by the GOCE mission. This is especially profound at the limits of the
141 spatial scales targeted by GOCE, i.e., around 100-120 km. For those levels that
142 the noise is either dominant or contaminates the gravity field signal, increasing the
143 SNR (signal-to-noise-ratio) demands a digital or spatial filter implementation. In
144 this work, both Gaussian and boxcar filters have been used to remove noise. The
145 former is considered very effective for filtering in the space domain and the latter
146 being a smoothing mathematical function, which uses a rectangular window in the
147 frequency domain. After this synthesis and filtering process, the new combined
148 GGMs can be evaluated w.r.t. some external information, e.g., GPS/Leveling
149 geoid heights and gravity anomalies, to quantify the improvement reached.

150

151 Another valuable tool in the evaluation of the GOCE GGMs is in terms of the
152 relationship between the GGM-derived gravity information and land topography.
153 Therefore, a further investigation in terms of the spectral coherence and
154 correlation was realized. Spectral coherence is a measure of the relation between
155 two signals or data sets and if applied to the analysis of gravity field data it can
156 give insight to how well they relate to the Earth's topography. The basic idea
157 behind this evaluation, performed for each level of WL decomposition, is that if a
158 GGM has higher coherence with the topography for a specific level, i.e., for a
159 specific spectral band, then it represents better compared to the other GGMs the
160 Earth's gravity field. In that way, the possible improvement by GOCE can be

161 envisaged at specific targeted wavebands. Coherence is defined as (Bendat and
162 Piersol, 2010):

$$C_{gH} = \frac{|G_{gH}|^2}{G_{gg}G_{HH}} \quad (2)$$

163
164 where, G_{gH} denotes the cross-spectral density between the gravity and topography
165 signals g and H , and G_{gg} and G_{HH} the auto-spectral density. Another useful
166 measure to be employed is the correlation matrix that can show whether and how
167 strongly two signals are related. In our case, we construct the correlation matrix
168 by estimating the correlation coefficients between the various levels of
169 decomposition of both gravity and topography. Then the correlation coefficients
170 are estimated as (Bendat and Piersol, 2010):

$$R_{gH}^{ij} = \frac{C_{gH}^{ij}}{\sqrt{C_{gg}^{ii}C_{HH}^{jj}}} \quad (3)$$

171
172 In Eq. (3), R_{gH}^{ij} denotes the correlation between level i of gravity data from the
173 GGMs and level j of the topography, C_{gH}^{ij} is the cross-covariance matrix of the
174 two datasets and C_{gg}^{ii} and C_{HH}^{jj} the covariance matrices. In that way, the correlation
175 matrix has as diagonal elements the coefficients for the same levels of
176 decomposition. Finally, it should be noted that we have used the “awkward” term
177 gravity to describe the GGM contribution in the sense that the gravity signal g can
178 be any functional related to the Earth’s gravity field. In our study, as far as
179 coherence and correlation are concerned, we focus on GGM-derived gravity
180 anomalies over the Amazon area and over Europe.

181

182 **2.2 Available data and external validation**

183 The present study focuses on the GGM evaluation with external data, over the
184 European Continent, within the region bounded between $30^\circ \leq \varphi \leq$

185 60° and $-10^\circ \leq \lambda \leq 30^\circ$. For the investigation of the correlation and coherence
186 between the topography and GGM-induced gravity, the study focuses on two
187 regions. The first one is the aforementioned European area while the second one
188 focuses over the Amazon region, bounded between $-10^\circ \leq \varphi \leq 20^\circ$ and $275^\circ \leq$
189 $\lambda \leq 335^\circ$. The main reason for focusing in two areas is their different topographic
190 features and gravity field representation. The former stems from the Amazon area
191 being mostly flat, while Europe has highly varying terrain. The latter refers to the
192 fact that Amazon is a poorly surveyed area in terms of terrestrial gravity data,
193 contrary to Europe. Therefore, GOCE GGMs should have improved performance
194 over the Amazon, in terms of the correlation and coherency with land topography
195 compared to EGM2008.

196

197 The GOCE/GRACE GGMs evaluated refer to the latest DIR and TIM models
198 employing the fourth (R4) and fifth (R5) release, i.e., an effective data volume of
199 26.5 months of GOCE observations (R4) and the low orbit data (R5). TIM-R4
200 (Pail et al., 2011) presents a spherical harmonics expansion of the Earth's
201 potential to a maximum degree $n_{max}=250$ while TIM-R5 reaches a maximum
202 degree of 280, both employing the time-wise approach and being pure GOCE
203 models. DIR-R4 (Bruinsma et al., 2010; 2013), presents a spherical harmonics
204 expansion of the Earth's potential to a maximum degree 260 while DIR-R5
205 reached a maximum degree 300, employing the direct approach and in addition to
206 the GOCE observations, 9 years of GRACE data (10 for DIR-R5) and SLR have
207 also been used. Apart from these GGMs, GOCO03s is used as well (Mayer-Gürr
208 et al. 2012), which is based on both GOCE and GRACE data with a maximum
209 degree of expansion to 250. In all cases, the performance of GOCE/GRACE
210 GGMs is evaluated against EGM2008 which is used as reference (Pavlis et al.,
211 2012). From these models, gravity anomalies and geoid heights have been
212 determined, for all areas under study, at grid points with a spatial resolution of
213 $3' \times 3'$ (5.5 km).

214

215 As far as the external GPS/Leveling and gravity data are concerned, the former
216 refer to a set of 1542 collocated GPS and Leveling observations over the Hellenic
217 trigonometric network (Vergos et al., 2014) in mainland Greece. The gravity data,
218 refer to the gravity anomaly field derived in the frame of the World Gravity Map

219 project (WGM2012) covering the entire European area (Balmino et al., 2012).
220 WGM2012 is largely based on EGM2008 with the addition of a worldwide 1'×1'
221 grid of ETOPO1-induced gravity disturbances. Therefore, the improvements
222 brought by GOCE to the gravity anomaly comparison are expected to be
223 marginal. Finally, the coherence and correlation between land topography and the
224 Earth's gravity field are computed for the Amazon and Europe area, where
225 ETOPO1 (Amante and Eakins, 2009) was used for the topographic information.
226 All computations were carried out in the Tide Free (TF) system, while GRS80 has
227 been used as a normal field in the evaluation of the geoid zero-degree term.
228

229 **3. GGM external validation with MRA**

230 Table 1 and Table 2 summarize the statistics of the differences between the
231 GPS/Levelling geoid heights, the WGM2010 gravity anomalies and the
232 GOCE/GRACE GGMs. In these Tables, the normal faced lettering denotes the
233 differences before any analysis of the GGM data has been performed, i.e., the
234 original ones to the GGM n_{max} . EGM2008, which will be used as the reference
235 GGM, provides a std at the 13.4 cm and 3.2 mGal, while the GOCE/GRACE
236 GGMs reach the 44-46 cm and 21-22 mGal, respectively. The latter is expected
237 due to the omission error in both geoid heights and gravity anomalies, due to the
238 small n_{max} .

239

240 The goal now is to try and enhance this performance of the GOCE/GRACE
241 GGMs through WL-based MRA. To that respect the derived geoid heights and
242 gravity anomalies from all GGMs have been decomposed through a discrete
243 wavelet transform (DWT) into 12 levels (L_1, L_2, \dots, L_{12}), where each level was
244 analyzed in an approximation coefficient and three detail coefficients. For all
245 GGMs, L_1 corresponds to spatial scales between 5.5-11 km (smallest scales), L_2 to
246 11-22 km, L_3 to 22-44 km, L_4 to 44-88 km, L_5 to 88-176 km, L_6 to 176-352 km, L_7
247 to 352-704 km, L_8 to 704-1,408 km, L_9 to 1,408-2,816 km, L_{10} to 2,816-5,632 km,
248 L_{11} to 5,632-11,264 km, and L_{12} to 11,264-22,528 km (largest scales). Given the
249 decomposition, the signals are then synthesized for the GOCE/GRACE GGMs, by
250 replacing some of their levels with those of EGM2008, so that a new combined
251 GGM would be generated, where both GOCE/GRACE and EGM2008

252 information is used. When the new GGM is denoted as, e.g., *GOCO03s* (L_5 , L_6 ,
253 L_7) this means that the first four levels come from EGM2008, levels 5-7 from
254 *GOCO03s* and then levels 8-12 from EGM2008 again. L_5 , L_6 , L_7 span the spatial
255 scales between 88 km and 704, corresponding to harmonic degrees ~ 28 -225,
256 therefore they represent the main waveband that GOCE should primarily provide
257 its highest improvement.

258

259 For the WL-based MRA implementation, various mother WLs have been tested,
260 from the simple Haar WL, to Coiflet and Daubechies (db), while for the latter two
261 various orders have been investigated. It was finally decided that db10, i.e., the
262 daubechies WL with ten vanishing moments, would be used for the WL analysis,
263 since it provided the best results to the GPS/Leveling geoid heights. After the
264 decomposition of the GGMs followed the reconstruction of their levels, by
265 combining their detail coefficients, and then the synthesis, as outlined in the
266 preceding section and Eq.1. In Tables 1 and 2 we summarize the synthesized
267 results from two scenarios, where GOCE/GRACE provides the information for L_5 ,
268 L_6 , L_7 and L_6 , L_7 . In the first case, the synthesized GGMs (herein denoted as
269 SynthGOCO03s, SynthTIM-R4, SynthTIM-R5, SynthDIR-R4 and SynthDIR-R5)
270 provide improved differences with the GPS/Leveling data by ~ 20 cm in terms of
271 the std compared to the original ones. SynthGOCO03s reduced the std to 26 cm,
272 while SynthTIM-R4/SynthTIM-R5 and SynthDIR-R4/SynthDIR-R5 are at the 24
273 cm and 22 cm. Still, these are about 10 cm worse than EGM2008, signalling that
274 some of the GOCE levels used for the synthesis are of lower quality and contain
275 more noise than signal. When only L_6 , L_7 from GOCE are used (corresponding to
276 spatial scales 176-352 km and 352-704 km, respectively) then the situation
277 improves drastically. SynthGOCO03s now reaches the 12.4 cm when compared to
278 the GPS/Leveling geoid heights, while SynthTIM-R4 and SynthDIR-R4 provide
279 even better results at the 12.2 and 12 cm. The R5 versions of the TIM and DIR
280 models manage to improve the results of their R4 counterparts to 12.1 and 11.8
281 cm. These are 1-1.6 cm better than the performance of EGM2008, signalling the
282 improvement offered by GOCE in the specific spectral range. Moreover, it shows
283 that the WL-based MRA is an effective tool in order to analyse the GOCE GGMs
284 and model their omission error in the shorter spatial scales.

285

286 **TABLE 1**

287

288 The same holds for the external evaluation with gravity anomalies, since the
289 Wavelet MRA Synthesis with L_5 , L_6 , L_7 improves the original ones by as much as
290 13-14 mGal. What is striking in the gravity anomaly differences is the significant
291 improvement brought by TIM-R5 and DIR-R5. The std of the differences before
292 any analysis is ~ 3 mGal lower than that of the R4 GGMs, while after the synthesis
293 they give improved std by ~ 1 mGal. This signals the value of the low orbit GOCE
294 data, which is especially evident in the gravity anomalies compared to geoid
295 heights due to the larger spectral content of the former in the smaller scales. When
296 the synthesis is based only on L_6 , L_7 from GOCE the std of the differences is the
297 same as that of EGM2008 (~ 3 mGal). It is not unexpected that the synthesized
298 GOCE/GRACE GGMs do not manage to improve the results of EGM2008, since
299 most of the WGM2012 gravity data have been used in the compilation of
300 EGM2008.

301

302 **TABLE 2**

303

304 From these results, it becomes apparent that indeed GOCE manages to improve
305 the geoid and gravity field representation in the medium bands of the spectrum
306 and that the followed methodology manages to provide reasonable results and can
307 be employed in related studies where a synthesized geoid and/or gravity field is
308 needed. A further test performed, was to examine the behavior of L_5 from the
309 GOCE/GRACE GGMs, given that its inclusion in the synthesis worsens the
310 results. L_5 corresponds to spatial scales between 88-176 km, therefore given that
311 GOCE was to provide useful signal up to spatial scales of 100 it means that part of
312 the signal in L_5 is beyond the measuring waveband of GOCE (ESA, 1999). To that
313 respect, and in order to remove some of the noise present in L_5 of the decomposed
314 GOCE/GRACE GGM signal, a simple Gaussian and a boxcar type of filter have
315 been tested. Both will act as low-pass filters, where we intend to remove some of
316 the high-frequencies in the waveband between 88-176 km. Various cut-off
317 wavelengths between 90 km and 140 km, with an increment of 5 km have been
318 tested, and finally it was decided to keep the one corresponding to a spatial scale
319 of 120 km. For all these tests, after the filtering, the signals were synthesized and

320 comparisons with the GPS/Leveling and gravity data have been performed. The
321 cut-off wavelength of 120 km was the one that offered the best results in the
322 comparisons with the external data. For smaller wavelengths, the results
323 deteriorate due to the presence of noise, while for larger wavelengths signal was
324 removed along with noise deteriorating the results. Tables 3 and 4 presents the
325 statistics of the differences between the filtered and then synthesized
326 GOCE/GRACE GGMs (denoted as, e.g., *GOCO03s-f* where *f* stands for filtered)
327 with the GPS/Leveling and gravity data. For GOCO03s, the std of the differences
328 with the GPS/Leveling data reduces from 25.9 cm to 17.6 cm and 19.2 cm for the
329 Gaussian and boxcar filters. A smaller improvement is found for the TIM-R4 and
330 DIR-R4 models, for which the std decreases by ~4-5 cm. Noticeably though,
331 while the synthesis of TIM-R4 (L_5 , L_6 , L_7) is better than that of GOCO03s (23.9
332 cm compared to 25.9 cm), after the filtering with the Gaussian filter GOCO03s is
333 better by 1.6 cm. This may be due to the fact that the latter is based on fewer
334 GOCE data, hence the L_5 geoid signal is poorly modeled compared to TIM-R4. In
335 that way, filtering benefits more the GGM with the least amount of information in
336 that waveband, by removing the noise, while in TIM-R4 it removes not only the
337 noise but some useful geoid signal that is present. With TIM-R4 and DIR-R4
338 probably a more elaborate filtering process is needed, so that the noise and signal
339 can be better decomposed and separated. This is out of the scope of the present
340 work and is actually the field of future investigation. For DIR-R4, after filtering
341 L_5 the std drops to 18.7 and 21.2 cm, showing slightly better results than TIM-R4.
342 Once again, more interesting results from the filtering are acquired for the R5
343 GGMs, since for TIM-R5 the std drops to 16.3 and 18.0 cm and for DIR-R5 to
344 15.7 and 17.7 cm, for the Gaussian and boxcar filter respectively. This is a 3 cm
345 improvement compared to the R4 GGMs and is due to the fact that the R5 models
346 contain the low orbit GOCE data, hence the SNR is larger at L_5 . The same results
347 are acquired for the external validation with gravity anomalies, where an
348 improvement of 2-3 mGal is achieved, compared to the unfiltered synthesis, but
349 this is still worse than the synthesis where only L_6 and L_7 are used. Concluding on
350 the filtering process investigated, the results acquired are indeed improved
351 compared to no filtering L_5 at all, but in no case did we achieve the results when
352 using for the synthesis only L_6 and L_7 . If some useful signal is to be derived from

353 L_5 , so that the results will be further improved, then more elaborate filtering
354 and/or error modeling is needed to de-correlate signal from noise.

355

356 **TABLE 3**

357

358 **TABLE 4**

359

360 The final part of the GOCE/GRACE GGM evaluation was carried out over the
361 Amazon area and Europe by investigating the spectral coherence and correlation
362 between the GGM-derived gravity anomalies and land topography. To derive the
363 coherence between the GGM gravity anomalies and topography/bathymetry, WL
364 MRA has been used again so that both signals have been decomposed in twelve
365 levels and then the signal of each level has been reconstructed. Then, employing
366 Eq. 2 the spectral coherence has been evaluated for each level between the
367 topography signal and gravity anomalies from EGM2008, TIM-R4/R5, DIR-
368 R4/R5 and GOCO03s.

369

370 For Amazon, the same results with the external evaluation have been derived,
371 with the higher coherence found in L_6 and L_7 for the GOCE GGMs and lower
372 coherence for L_5 and L_4 . These are depicted in Figure 1, where the coherence for
373 L_3 , L_5 , L_6 and L_7 is shown for the area over Amazon. As expected, the coherence
374 for the GOCE/GRACE GGMs for L_3 is very low with practically no coherence up
375 to 30-35 km and then only up to 20%, which is probably just noise. EGM2008 on
376 the other hand has a more or less uniform coherence between 55% and 72%. In L_5
377 the situation starts to improve for the GOCE/GRACE GGMs, with higher
378 coherence up to ~42%, while in all cases their coherence is below that of
379 EGM2008. Notice that the filtering cut-off wavelength set in the previous test to
380 120 km coincides in the coherence plot with the point that the coherence starts to
381 raise for the GOCE/GRACE GGMs. Therefore, in the second half of the
382 GOCE/GRACE L_5 there seems indeed to be some useful signal that still remains
383 to be accounted for within the present methodological scheme. The situation
384 changes completely for L_6 and L_7 , where the GOCE /GRACE GGMs are
385 equivalent to EGM2008 and in most cases better than that. Between the
386 GOCE/GRACE GGMs it is interesting to notice that for L_5 , the release 5 versions

387 of TIM and DIR have better coherency than GOCO, which is due to the more
388 GOCE data used in their development. For larger wavelengths, GOCO performs
389 equally well given that these lower harmonic degrees of the spectrum are
390 sufficiently modeled by the release 5 of GOCE data along with the information
391 from GRACE which was used in its development. Comparing the R4 and R5
392 versions of the TIM and DIR models, Figure 2 shows their coherence for L_5 over
393 Amazon. From this Figure it is readily seen how the addition of the low-orbit
394 GOCE data manage to provide improved results by about 10-20%. This is
395 especially evident at spatial scales between 140-160 km, where the difference
396 between TIM-R4 and TIM-R5 is at the 20% level. Finally, DIR-R5 manages to
397 provide improved results especially for scales between 80-100 km, where it is
398 better by as much as 35%.

399

400 **Figure 1**

401

402

403 **Figure 2**

404

405 Over Europe the results are slightly different, given that EGM2008 is dominant.
406 For L_3 EGM2008 has a mean coherence of 70% while it reaches the 85% level as
407 well. For L_5 EGM2008 retains a high coherence between 50% and 80% while the
408 GOCE GGMs have a maximum of 50% and a mean of ~30%. This is due to fact
409 that Europe is a well surveyed area in terms of terrestrial gravity data, hence
410 EGM2008 manages to represent very well the gravity field over this region. From
411 this analysis it can be concluded that indeed in poorly surveyed areas, GOCE
412 GGMs can be expected to contribute significantly in gravity field mapping.

413

414 Likewise, the correlation between the topographic signal and that of the GGM-
415 derived gravity anomalies for the various levels of decomposition has been
416 analyzed. Table 5 presents the correlation coefficients for EGM2008 and the
417 GOCE/GRACE GGMs only for Amazon given the aforementioned discussion.
418 The correlation results for Europe are again in favor of EGM2008, since most of
419 the terrestrial gravity anomalies over the entire continent have been used in its
420 development. A similar picture with the coherence analysis is found. For the very

421 first levels (L_1 to L_4), EGM2008 dominates since the GOCE/GRACE GGMs
422 present little or no correlation at all (up to 33% for L_4). From L_5 onwards the
423 contribution of GOCE is evident, since the latter GGMs manage to perform
424 equally well with EGM2008 (for L_5) and outperform it for L_6 , L_7 and L_8 . For L_6
425 and L_7 they present a higher correlation between 3-5% compared to EGM2008,
426 which increases to 22% for L_8 . The correlation found for L_9 (corresponding to
427 spatial scales between 1408-2816km) is somewhat puzzling, since it is lower than
428 that for L_8 and L_{10} . Analyzing the topography and gravity anomaly signals for L_9 ,
429 it was found that the former has two (positive) dominant features in the EW
430 direction with a low over the Amazon basin, probably coming from the Andes to
431 the East and a merge of the Brazilian and Guyana shields in the West. On the
432 other hand, the gravity anomaly signal shows very little variation for these spatial
433 scales over the area under study, with two predominant positive features in the NS
434 direction and a low over the Amazon basin. This inconsistency can be due to the
435 fact that the gravity signal from the Andes is more high frequency in nature and is
436 represented in the lower levels (it is mostly seen in L_4 - L_7) compared to the
437 Brazilian and Guyana shield topography that dominate the area under study in the
438 southern and northern parts respectively. In any case, the contribution of GOCE to
439 the medium wavebands of the spectrum is once again evident, while if a higher
440 resolution digital terrain model have been used, then the superiority of the
441 GOCE/GRACE GGMs would have been more evident in the coherence and
442 correlation analysis.

443

444 **TABLE 5**

445

446 **4. Conclusions**

447 A detailed evaluation, employing WL-based MRA, has been carried out for the
448 recent GOCE and GOCE/GRACE GGMs both in terms of geoid heights and
449 gravity anomalies. From the external evaluation that referred to geoid heights, it
450 was concluded that the combined GGMs improve the estimated geoid heights,
451 compared to local GPS/Leveling data, since the std is reduced from ± 0.45 m to
452 ± 0.22 m for the DIR-R4/R5 and TIM-R4/R5 models. When only L_6 and L_7 have
453 been used from the GOCE/GRACE GGMs, then the results improved further to

454 ± 0.118 cm for DIR-R5, being 1.6 cm better than EGM2008. Contrary to that, the
455 evaluation with gravity anomalies revealed that in the best case scenario the
456 GOCE/GRACE GGMs manage to reach the agreement of EGM2008, something
457 attributed to the fact that most of the data used for the compilation of WGM2012
458 have been included in EGM2008. Hence this gravity information is not a totally
459 independent dataset due to the high correlation with EGM2008.

460

461 When filtering L_5 of the GOCE/GRACE GGMs the results improved between 4-8
462 cm and 2-3 mGal, showing that indeed L_5 contains useful geoid/gravity signal that
463 can contribute to the overall GGM performance. To that respect, more advanced
464 filtering options will be investigated to separate the noise from the signal and
465 improve the GGM performance. Overall, the proposed WL MRA methodology,
466 for the analysis and synthesis of GOCE/GRACE GGMs, provides promising
467 results since spatial scales in the GGMs that are modeled with lower accuracy can
468 be successfully replaced with other sources of information which are of higher
469 accuracy/quality. It is worth mentioning that with the presented WL-based
470 analysis there is no mixing of the spatial scales of the gravity field signal. This is
471 so in the sense that WLs can isolate specific portions of the gravity field signal
472 through the analysis in various levels of decomposition, the latter corresponding
473 to specific spatial scales. Then each level can be manipulated separately from the
474 rest, allowing the study of the gravity field signal properties for the specific spatial
475 scales. This cannot be done with, e.g., Fourier transform based methods, where
476 the entire spectrum of the gravity field signal is studied in its entity. From our
477 analysis, the main problem with this WL-based approach is that the dyadic nature
478 of WLs allow the isolation of specific spatial scales for each level (given the
479 resolution of the original signal). If one would like to study, e.g., the GOCE GGM
480 signal for spatial scales between 60 and 140 km only, then this approach cannot
481 be followed, since the specific range belongs to two different levels (L_4 and L_5) so
482 each level should be studied separately.

483

484 From the analysis of the spectral coherency and correlation between topography
485 and the GGM-derived gravity anomalies it was concluded that EGM2008 has
486 significantly better results for the first few levels. This is expected since the
487 GOCE/GRACE GGMs investigated are satellite-only ones. The contribution of

488 the satellite missions in seen again in L_5 , L_6 , L_7 and L_8 , where the GOCE/GRACE
489 GGMs show improved result compared to EGM2008 by 3-22%.

490

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492

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Table 1: Statistics of the differences between GPS/levelling and geoid heights from the GGMs before (normal font) and after (italics) the WL MRA synthesis. Units: [m]

	min	max	mean	std
EGM08	-0.853	0.104	-0.372	±0.134
GOCO03s	-1.735	1.110	-0.359	±0.464
<i>GOCO03s (L₅,L₆, L₇)</i>	<i>-1.083</i>	<i>0.453</i>	<i>-0.387</i>	<i>±0.259</i>
<i>GOCO03s (L₆, L₇)</i>	<i>-0.855</i>	<i>0.093</i>	<i>-0.378</i>	<i>±0.124</i>
TIM-R4	-1.597	1.155	-0.358	±0.450
<i>TIM-R4 (L₅,L₆, L₇)</i>	<i>-1.151</i>	<i>0.399</i>	<i>-0.381</i>	<i>±0.239</i>
<i>TIM-R4 (L₆, L₇)</i>	<i>-0.838</i>	<i>0.053</i>	<i>-0.383</i>	<i>±0.122</i>
TIM-R5	-1.569	1.123	-0.394	±0.469
<i>TIM-R5 (L₅,L₆, L₇)</i>	<i>-1.142</i>	<i>0.408</i>	<i>-0.393</i>	<i>±0.242</i>
<i>TIM-R5 (L₆, L₇)</i>	<i>-0.831</i>	<i>0.047</i>	<i>-0.383</i>	<i>±0.121</i>
DIR-R4	-1.540	1.105	-0.366	±0.442
<i>DIR-R4 (L₅,L₆, L₇)</i>	<i>-1.048</i>	<i>0.401</i>	<i>-0.392</i>	<i>±0.223</i>
<i>DIR-R4 (L₆, L₇)</i>	<i>-0.802</i>	<i>0.064</i>	<i>-0.394</i>	<i>±0.120</i>
DIR-R5	-1.530	1.122	-0.388	±0.454
<i>DIR-R5 (L₅,L₆, L₇)</i>	<i>-1.031</i>	<i>0.388</i>	<i>-0.389</i>	<i>±0.217</i>
<i>DIR-R5 (L₆, L₇)</i>	<i>-0.811</i>	<i>0.032</i>	<i>-0.381</i>	<i>±0.118</i>

Table 2: Statistics of the gravity anomaly differences between WGM2012 and the GGMs before (normal font) and after (italics) the WL MRA synthesis. Units: [mGal]

	min	max	mean	std
EGM08	-49.66	128.50	0.31	±3.24
GOCO03s	-204.97	272.23	0.11	±22.49
<i>GOCO03s (L₅,L₆, L₇)</i>	<i>-89.09</i>	<i>129.87</i>	<i>0.32</i>	<i>±9.38</i>
<i>GOCO03s (L₆, L₇)</i>	<i>-51.19</i>	<i>123.95</i>	<i>0.31</i>	<i>±3.52</i>
TIM-R4	-206.98	269.35	0.11	±22.14
<i>TIM-R4 (L₅,L₆, L₇)</i>	<i>-90.52</i>	<i>134.80</i>	<i>0.31</i>	<i>±8.85</i>
<i>TIM-R4 (L₆, L₇)</i>	<i>-50.17</i>	<i>123.54</i>	<i>0.30</i>	<i>±3.48</i>
TIM-R5	-196.90	272.70	0.11	±19.34
<i>TIM-R5 (L₅,L₆, L₇)</i>	<i>-59.78</i>	<i>129.77</i>	<i>0.49</i>	<i>±7.66</i>
<i>TIM-R5 (L₆, L₇)</i>	<i>-40.53</i>	<i>124.22</i>	<i>0.43</i>	<i>±3.37</i>
DIR-R4	-201.93	271.43	0.11	±21.93
<i>DIR-R4 (L₅,L₆, L₇)</i>	<i>-87.10</i>	<i>129.69</i>	<i>0.29</i>	<i>±8.47</i>
<i>DIR-R4 (L₆, L₇)</i>	<i>-51.12</i>	<i>123.89</i>	<i>0.28</i>	<i>±3.44</i>
DIR-R5	-203.66	270.65	0.11	±19.10
<i>DIR-R5 (L₅,L₆, L₇)</i>	<i>-65.17</i>	<i>129.42</i>	<i>0.54</i>	<i>±7.38</i>
<i>DIR-R5 (L₆, L₇)</i>	<i>-41.02</i>	<i>125.50</i>	<i>0.41</i>	<i>±3.35</i>

Table 3: Statistics of the differences between GPS/levelling and geoid heights from the synthesized GGMs after filtering GOCE/GRACE L₅. Normal lettering for Gauss and italics for boxcar filtering. Units: [m]

	min	max	mean	std
GOCO03s-f (L ₅ ,L ₆ , L ₇)	-0.870	0.206	-0.377	±0.176
<i>GOCO03s-f (L₅,L₆, L₇)</i>	<i>-0.898</i>	<i>0.244</i>	<i>-0.373</i>	<i>±0.192</i>
TIM-R4-f (L ₅ ,L ₆ , L ₇)	-0.927	0.279	-0.377	±0.190
<i>TIM-R4-f (L₅,L₆, L₇)</i>	<i>-1.011</i>	<i>0.314</i>	<i>-0.373</i>	<i>±0.213</i>
TIM-R5-f (L ₅ ,L ₆ , L ₇)	-0.907	0.202	-0.384	±0.163
<i>TIM-R5-f (L₅,L₆, L₇)</i>	<i>-0.874</i>	<i>0.237</i>	<i>-0.379</i>	<i>±0.180</i>
DIR-R4-f (L ₅ ,L ₆ , L ₇)	-0.925	0.279	-0.378	±0.187
<i>DIR-R4-f (L₅,L₆, L₇)</i>	<i>-1.027</i>	<i>0.316</i>	<i>-0.373</i>	<i>±0.212</i>
DIR-R5-f (L ₅ ,L ₆ , L ₇)	-0.875	0.189	-0.381	±0.157
<i>DIR-R5-f (L₅,L₆, L₇)</i>	<i>-0.874</i>	<i>0.228</i>	<i>-0.376</i>	<i>±0.177</i>

Table 4: Statistics of the differences between WGM2012 and gravity anomalies from the synthesized GGMs after filtering GOCE/GRACE L_5 . Normal lettering for Gauss and italics for boxcar filtering. Units: [mGal]

	min	max	mean	std
GOCO03s-f (L_5, L_6, L_7)	-72.77	127.91	0.31	±6.48
<i>GOCO03s-f</i> (L_5, L_6, L_7)	-74.65	130.64	0.32	±6.94
TIM-R4-f (L_5, L_6, L_7)	-74.88	132.14	0.29	±6.36
<i>TIM-R4-f</i> (L_5, L_6, L_7)	-75.29	135.47	0.29	±6.81
TIM-R5-f (L_5, L_6, L_7)	-51.01	130.51	1.09	±5.73
<i>TIM-R5-f</i> (L_5, L_6, L_7)	-53.43	134.46	1.35	±6.19
DIR-R4-f (L_5, L_6, L_7)	-73.01	130.56	0.29	±6.25
<i>DIR-R4-f</i> (L_5, L_6, L_7)	-73.96	134.30	0.29	±6.72
DIR-R5-f (L_5, L_6, L_7)	-53.83	129.66	1.09	±5.63
<i>DIR-R5-f</i> (L_5, L_6, L_7)	-55.70	132.90	1.35	±6.14

Table 5: Correlation between Topography and GGM-derived gravity anomalies over Amazon.

	EGM2008	GOCO03s	TIM-R4	TIM-R5	DIR-R4	DIR-R5
L₁	28.80%	0.10%	2.10%	1.80%	0.30%	0.50%
L₂	59.80%	0.20%	1.40%	0.90%	0.20%	0.21%
L₃	71.90%	3.30%	5.50%	8.58%	2.50%	8.27%
L₄	74.20%	28.20%	32.10%	31.70%	30.10%	33.47%
L₅	65.50%	60.80%	65.20%	63.90%	63.00%	63.90%
L₆	62.40%	63.10%	66.20%	69.87%	68.70%	69.87%
L₇	67.70%	70.20%	71.50%	74.10%	70.10%	74.07%
L₈	42.10%	44.50%	46.20%	64.30%	45.10%	64.16%
L₉	23.00%	29.20%	25.50%	26.00%	26.50%	29.80%
L₁₀	30.10%	35.00%	39.10%	53.20%	39.80%	53.60%
L₁₁	51.80%	52.10%	46.80%	56.10%	52.30%	64.20%
L₁₂	97.90%	98.10%	98.20%	99.00%	98.50%	99.63%

Figure 1: Spectral coherency for various levels of decomposition between topography and GGM-derived gravity anomalies over the Amazon area.

Figure 2: Spectral coherency for L_5 between topography and TIM-R4/R5 and DIR-R4/R5 GGM-derived gravity anomalies over the Amazon area.